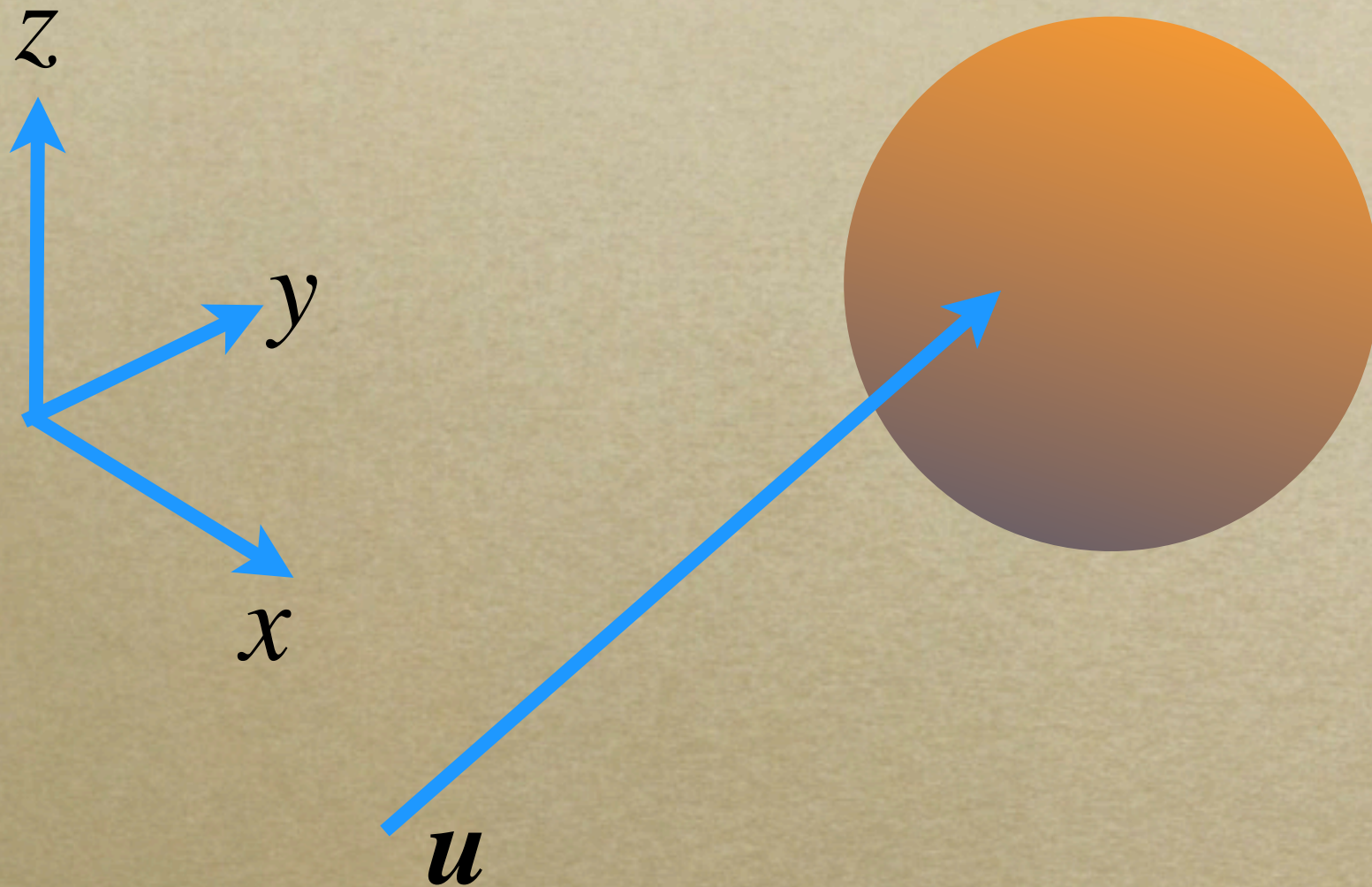
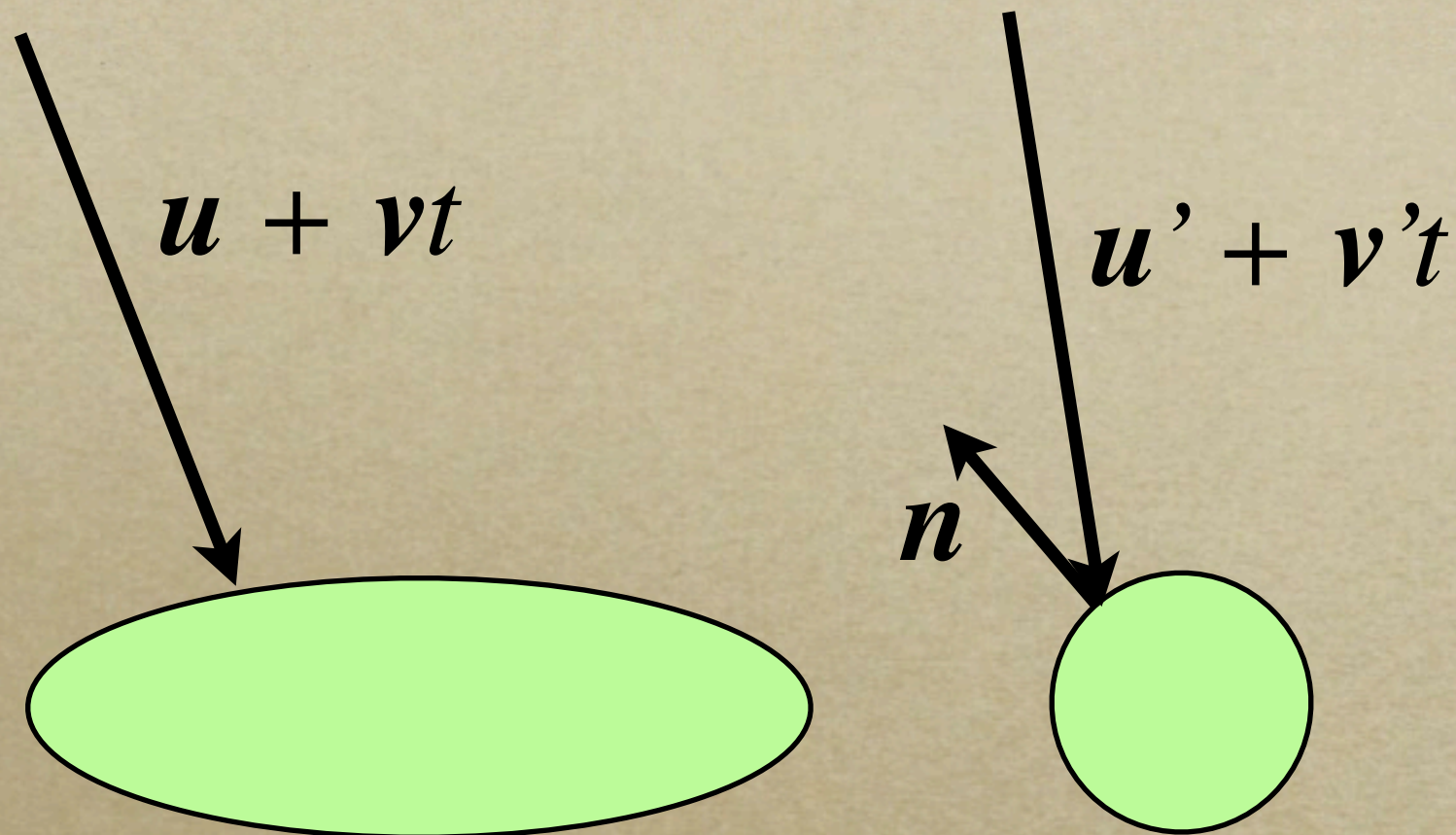


*Efficient illumination of
transformed objects
and some global illumination*

Transformed objects

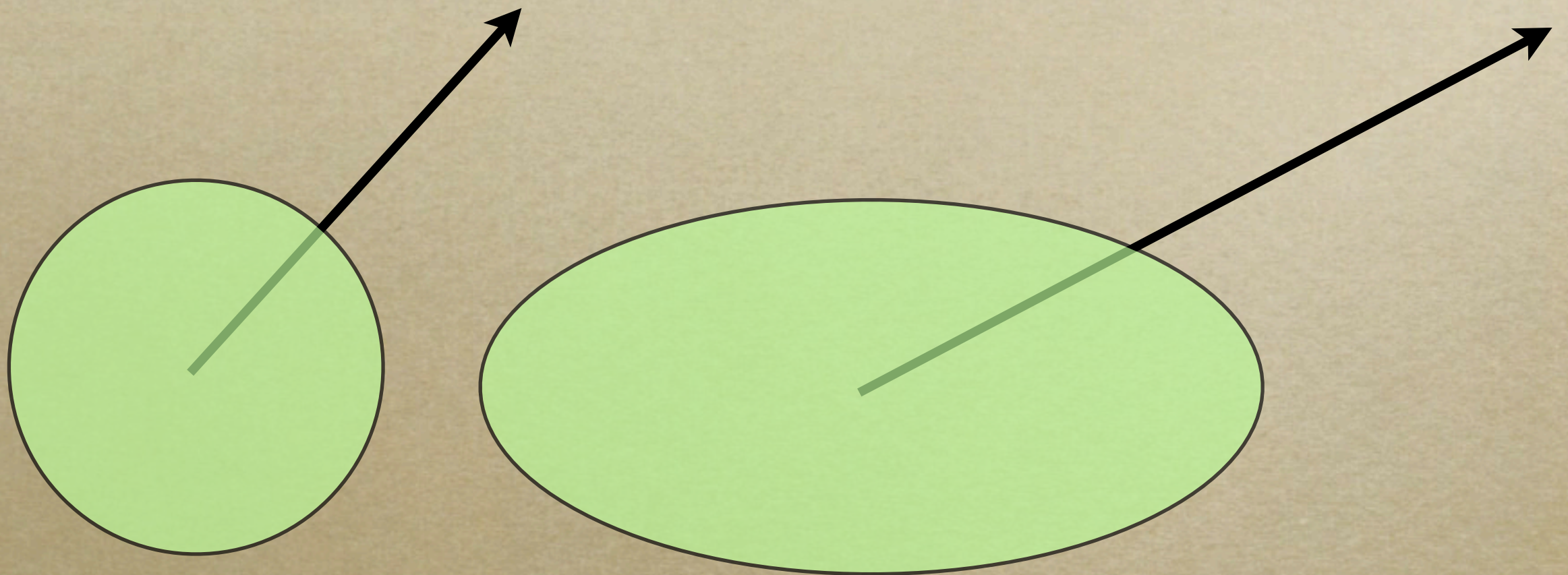


Transform the ray but ...



How do you get the surface normal back into world space?

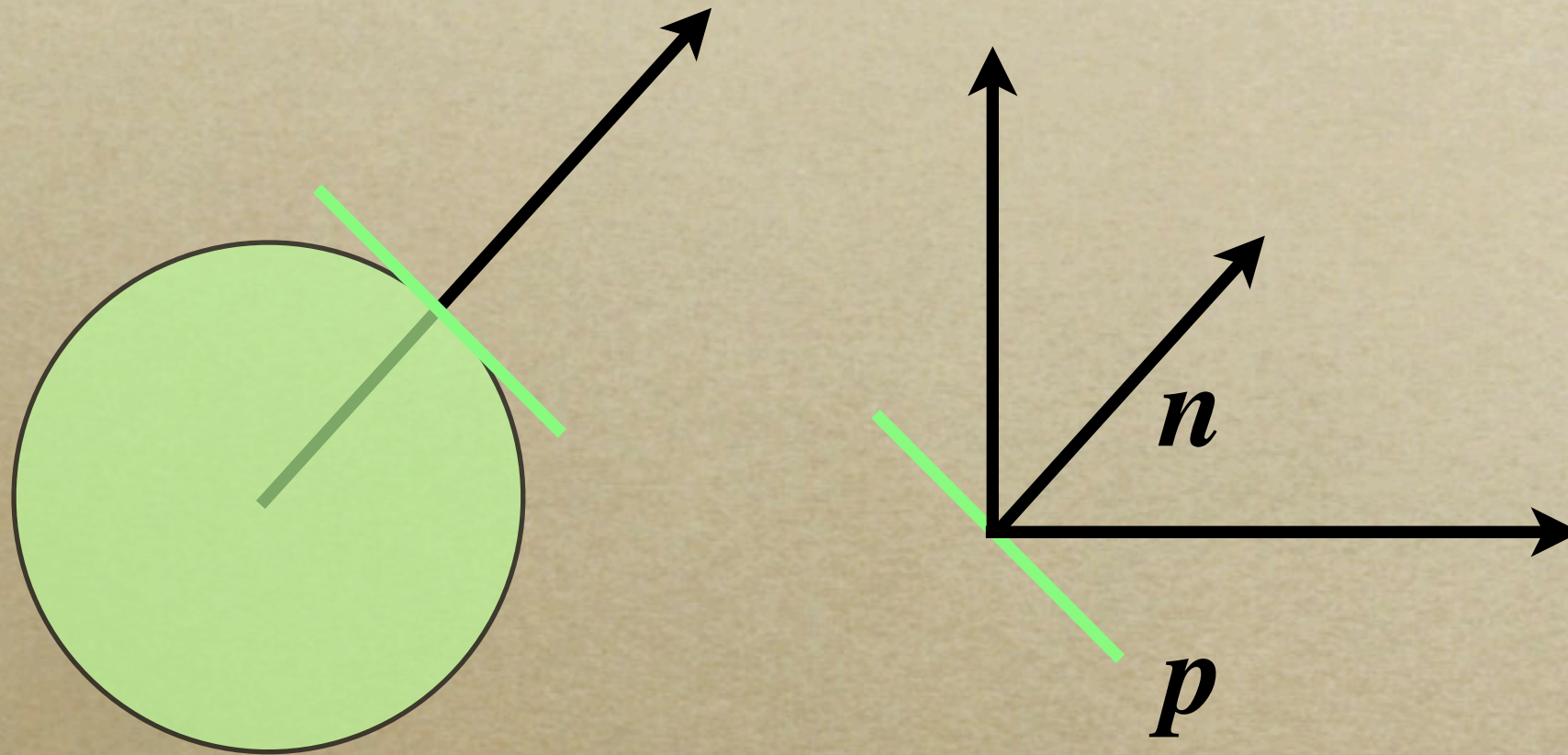
Transformed n is not normal



So what do we do?

- *When we apply a transformation matrix lines and planes are preserved but not angles.*
- *The normal defines a plane and the plane transforms to a plane.*

Which Plane?



*$p \cdot n = 0$ is a plane through the origin.
Suppose $p' = Tp$*

Dot product as matrix multiplication

$$\mathbf{p} = \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix} \quad \mathbf{v} = \begin{bmatrix} a \\ b \\ c \\ 0 \end{bmatrix}$$

$$\mathbf{p}^T \mathbf{v} = \begin{bmatrix} x & y & z & 1 \end{bmatrix} \begin{bmatrix} a \\ b \\ c \\ 0 \end{bmatrix} = \mathbf{p} \cdot \mathbf{v}$$

Now we can do the maths

$\mathbf{p} \cdot \mathbf{n} = 0$ is a plane through the origin.

or in matrix terms $\mathbf{p}^T \mathbf{n} = 0$

Suppose $\mathbf{p}' = T\mathbf{p}$, (T is any transformation)

$\mathbf{p} = T^{-1}\mathbf{p}'$ and $T^{-1}\mathbf{p}' \cdot \mathbf{n} = 0$

So... $T^{-1}\mathbf{p}' \cdot \mathbf{n} = 0$ or $(T^{-1}\mathbf{p}')^T \mathbf{n} = 0$

Now $(AB)^T = B^T A^T$ (prove that yourself)

$(\mathbf{p}'^T T^{-1T}) \mathbf{n} = 0$ $\mathbf{p}'^T (T^{-1T} \mathbf{n}) = 0$

i.e.: $\mathbf{p}' \cdot T^{-1T} \mathbf{n} = 0$

Beyond maths: see what it means

$$\mathbf{p} \cdot \mathbf{n} = 0 \quad \text{and} \quad \mathbf{p}' = T\mathbf{p}$$

\mathbf{p} is a point on a plane with normal \mathbf{n} .

\mathbf{p}' is a point on a transformed plane.

And we have shown that $\mathbf{p}' \cdot (T^{-1})^T \mathbf{n} = 0$

So \mathbf{p}' is a point on a plane with normal

$$T^{-1T} \mathbf{n}$$

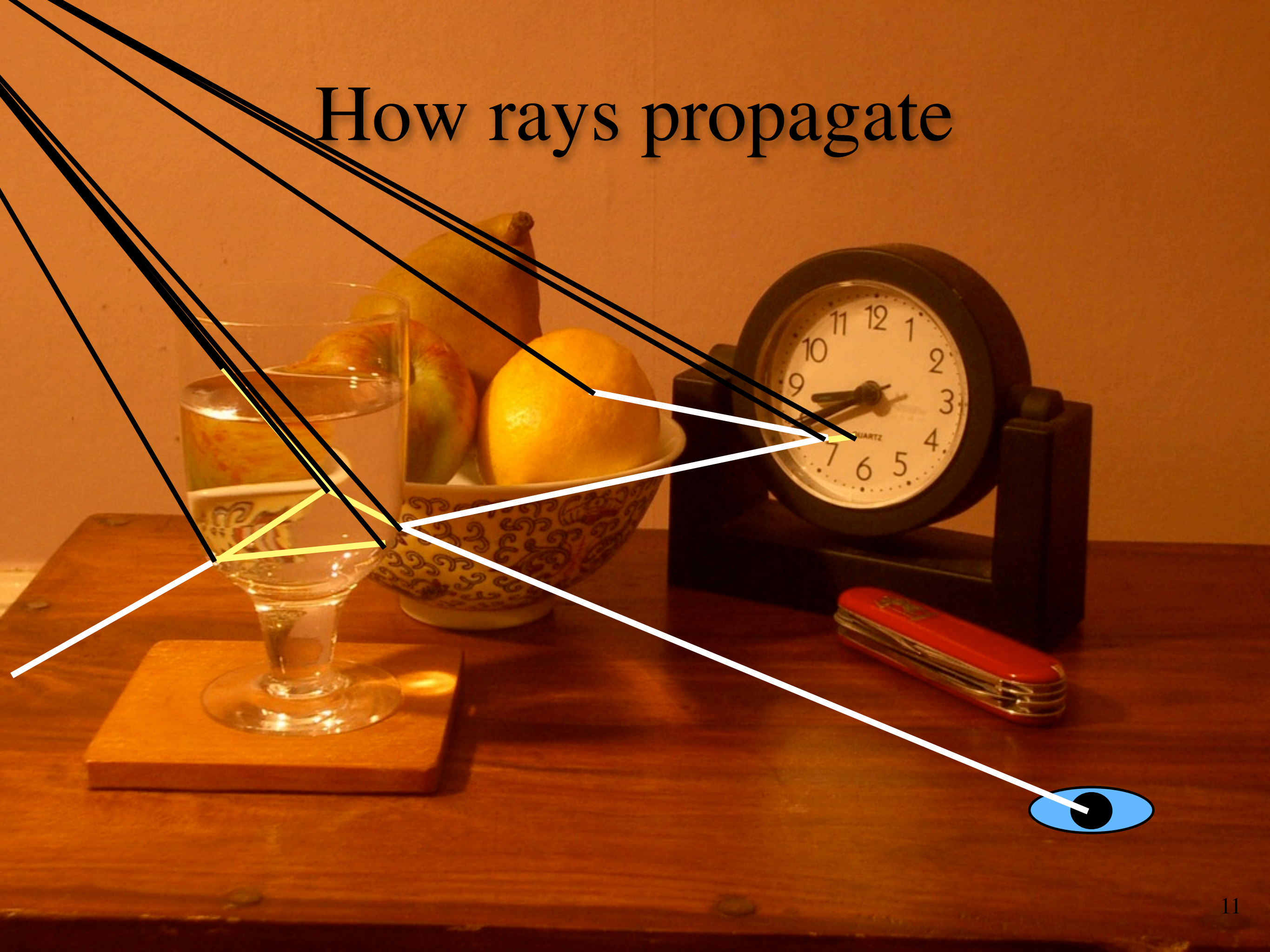
And the application...

Suppose we have an object that has been transformed by a matrix T .

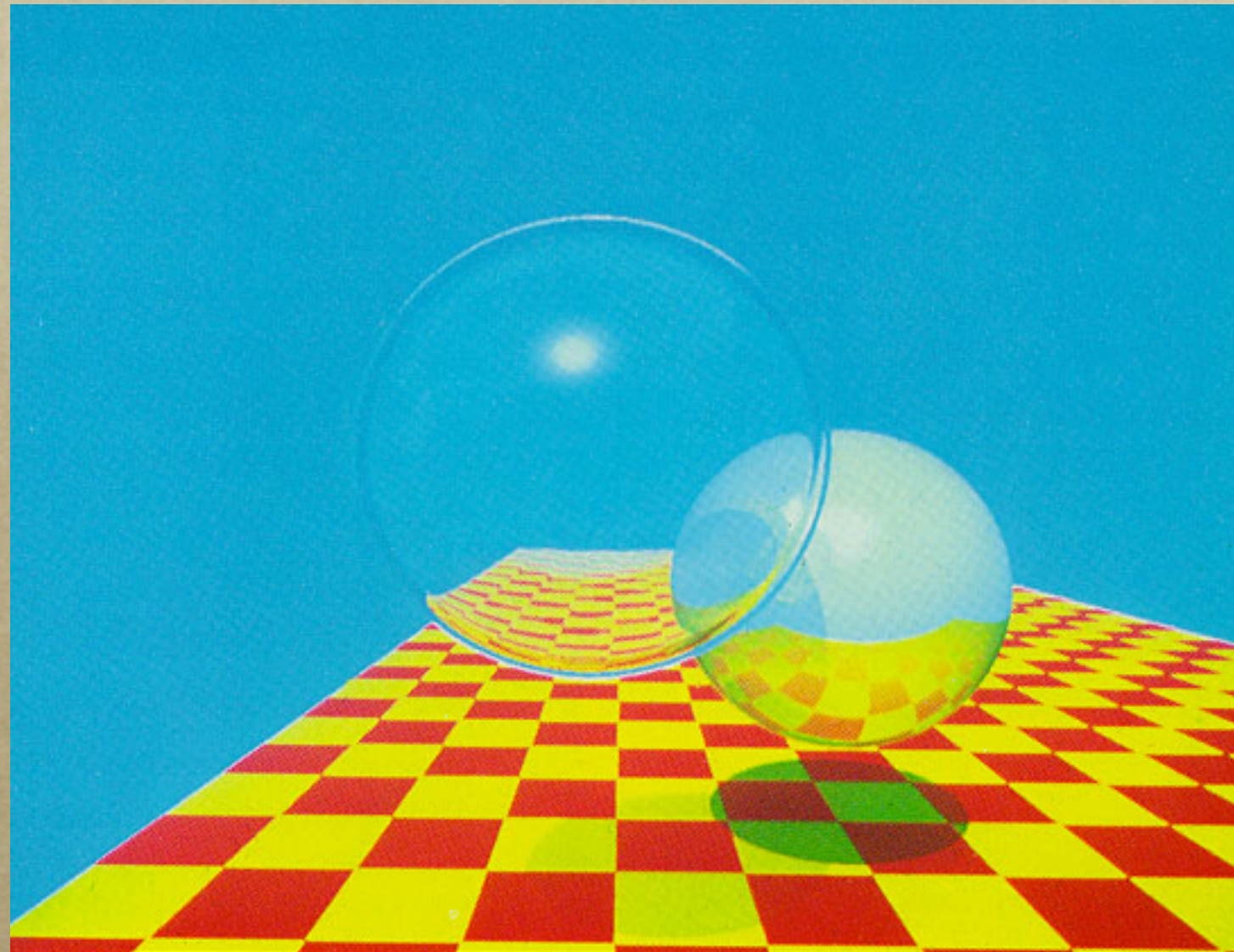
We transform $\mathbf{u} + \mathbf{v}t$ by T^{-1} and find t at the hit point, and a normal, \mathbf{n} .

The hit point in world space is $\mathbf{u} + \mathbf{v}t$ and the normal in world space is $T^{-1T}\mathbf{n}$.

How rays propagate



Whitted 1980



*Ambient, Lambert, Phong, reflection,
refraction, point light sources.*

Just the beginning...

- *Aliasing artefacts*
- *No surface/surface illumination*
- *No caustics*
- *Real shadows are soft*
- *Colour problems*
- *Very slow*

Just the beginning...

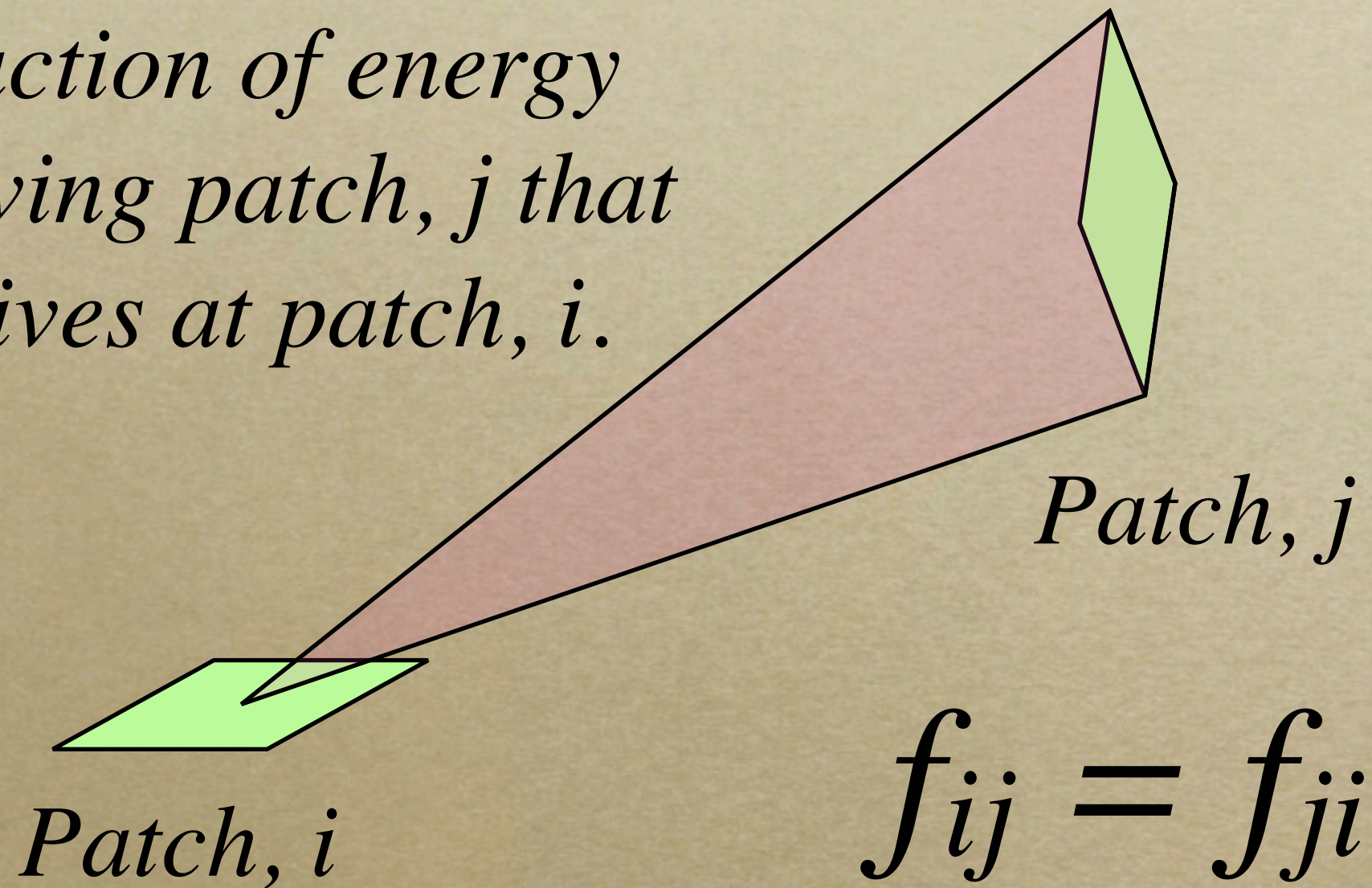
- *Aliasing artefacts*
- ***No surface/surface illumination***
- *No caustics*
- *Real shadows are soft*
- *Colour problems*
- *Very slow*

Radiosity

- *Divide the scene into small surface patches.*
- *For every patch pair find form factor.*
- *Find radiosities*
- *Render picture*

Form Factor, $f_{i,j}$

*Fraction of energy
leaving patch, j that
arrives at patch, i .*

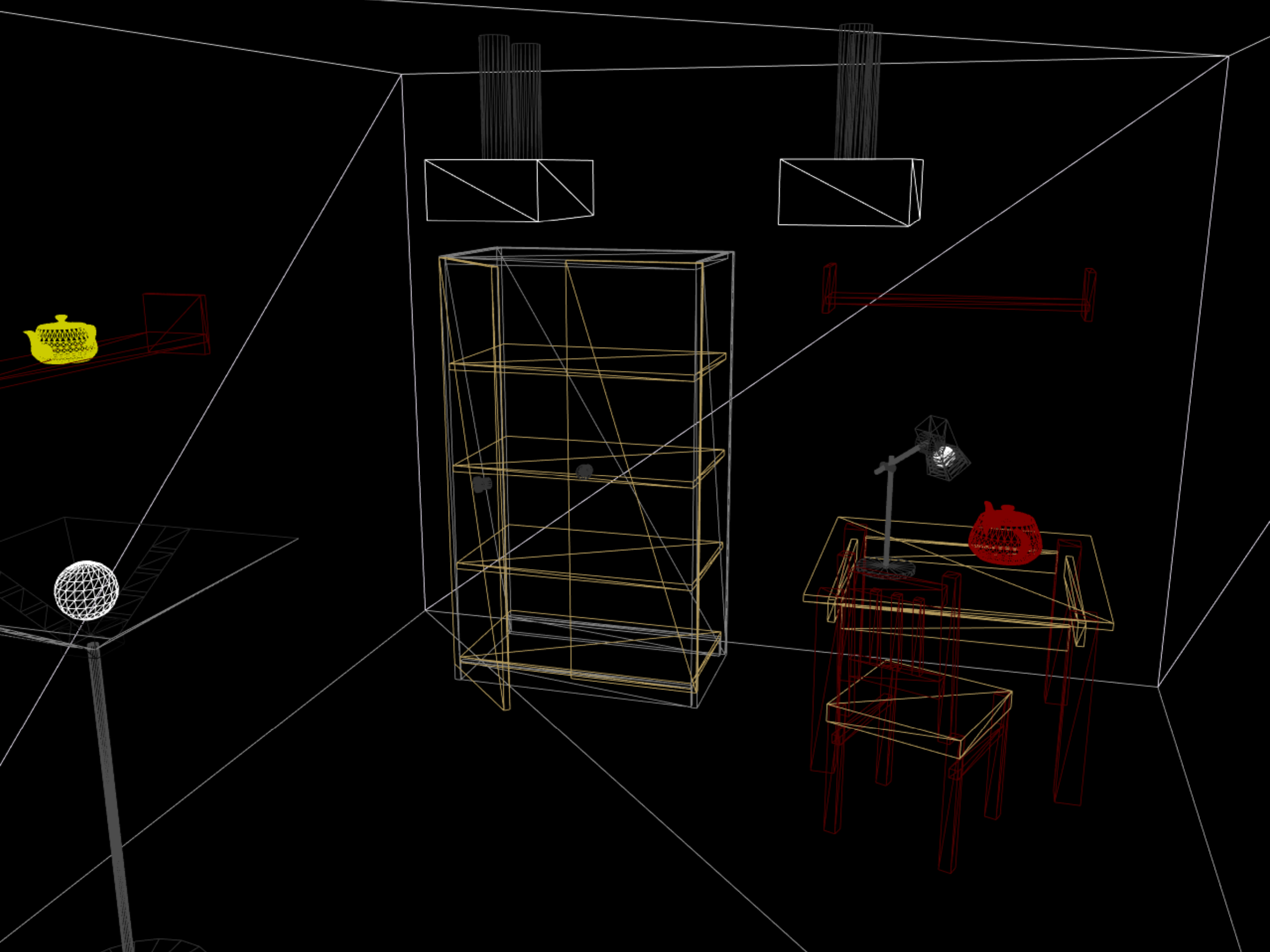


Radiosity Equation

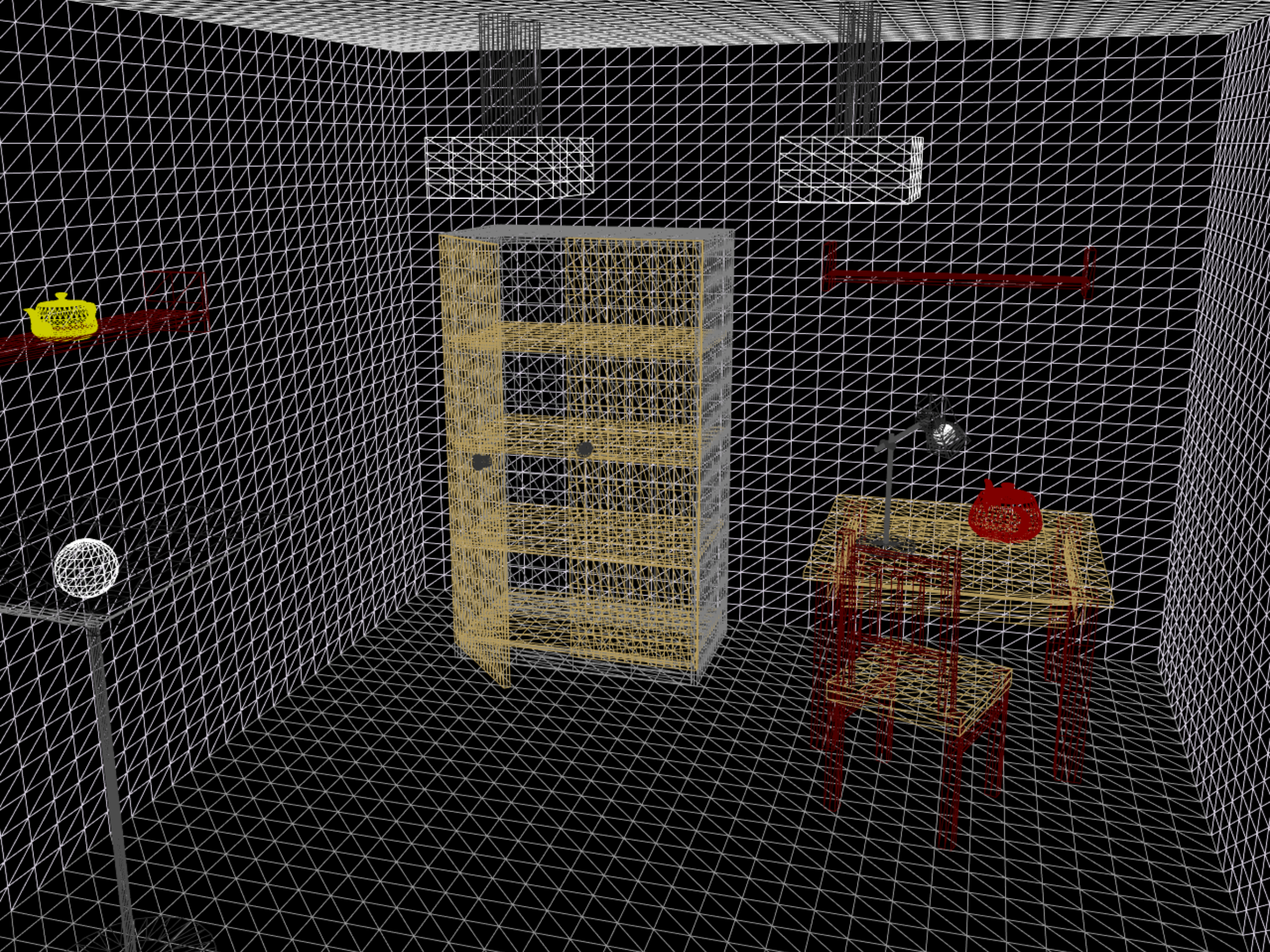
$$R_i = E_i + k_{d,i}(\sum f_{i,j}R_j)$$

Set up this system of equations and solve.

Or use successive approximation.

















Just the beginning...

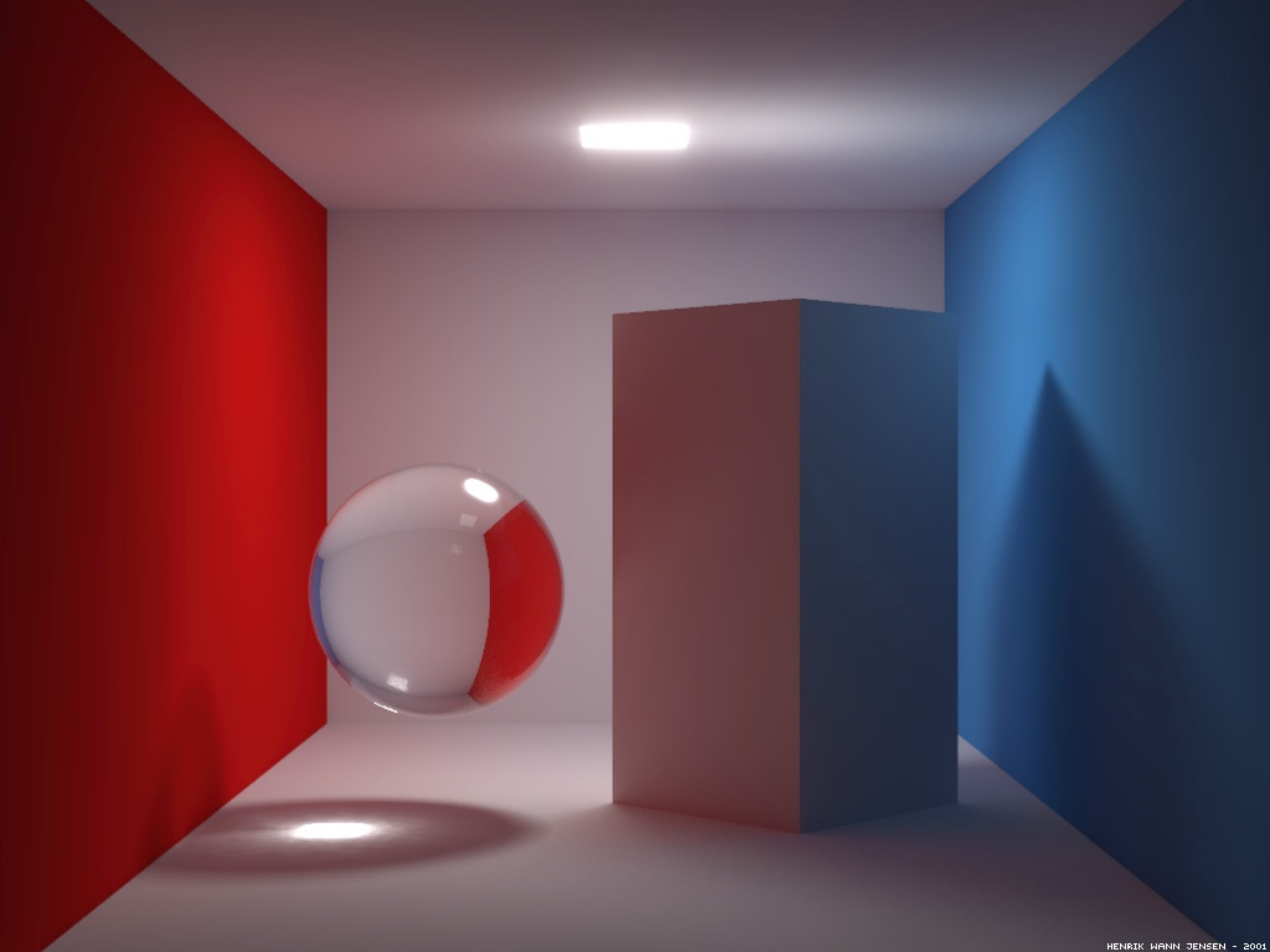
- *Aliasing artefacts*
- *No surface/surface illumination*
- ***No caustics***
- *Real shadows are soft*
- *Colour problems*
- *Very slow*

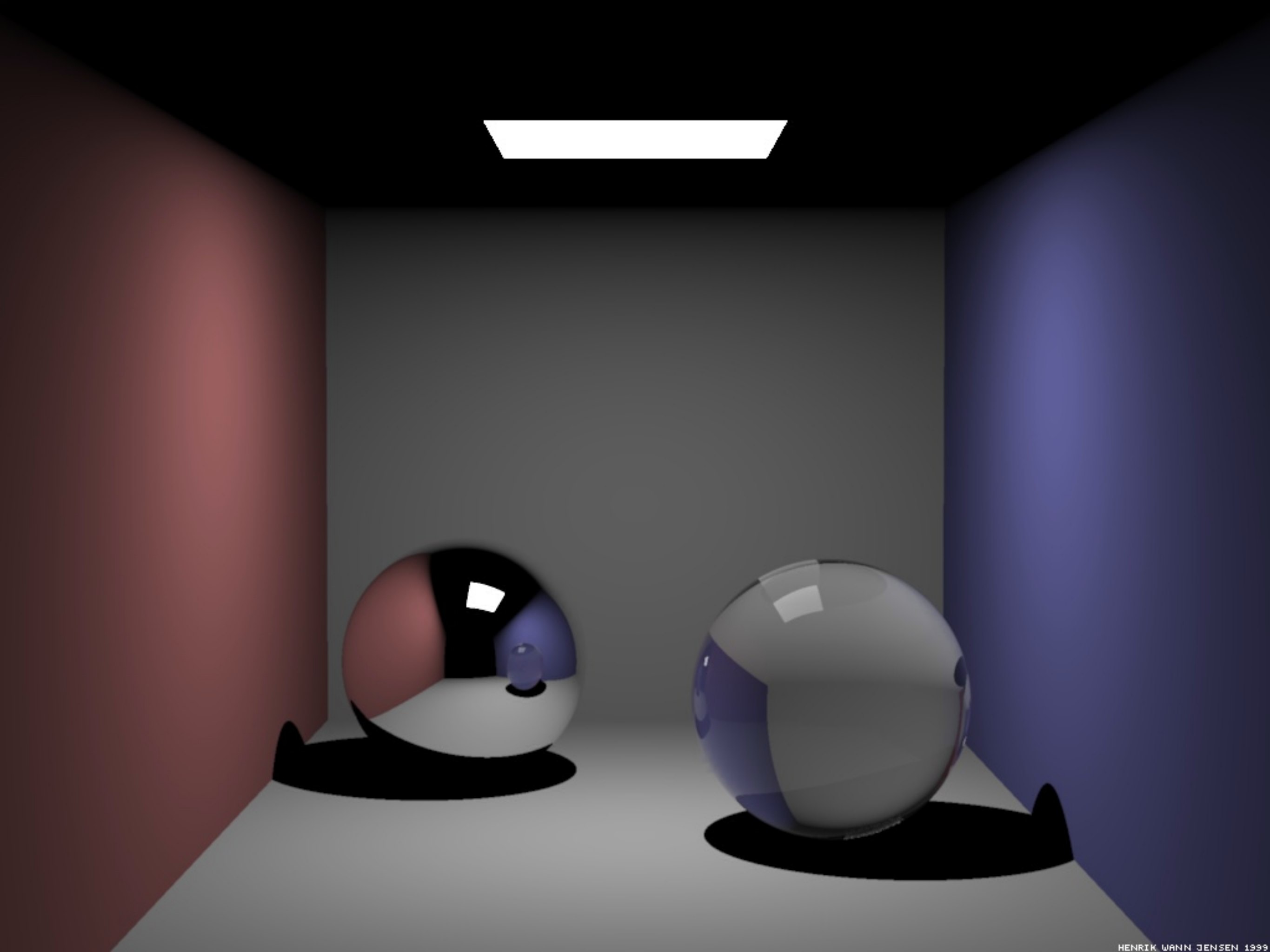
Caustics

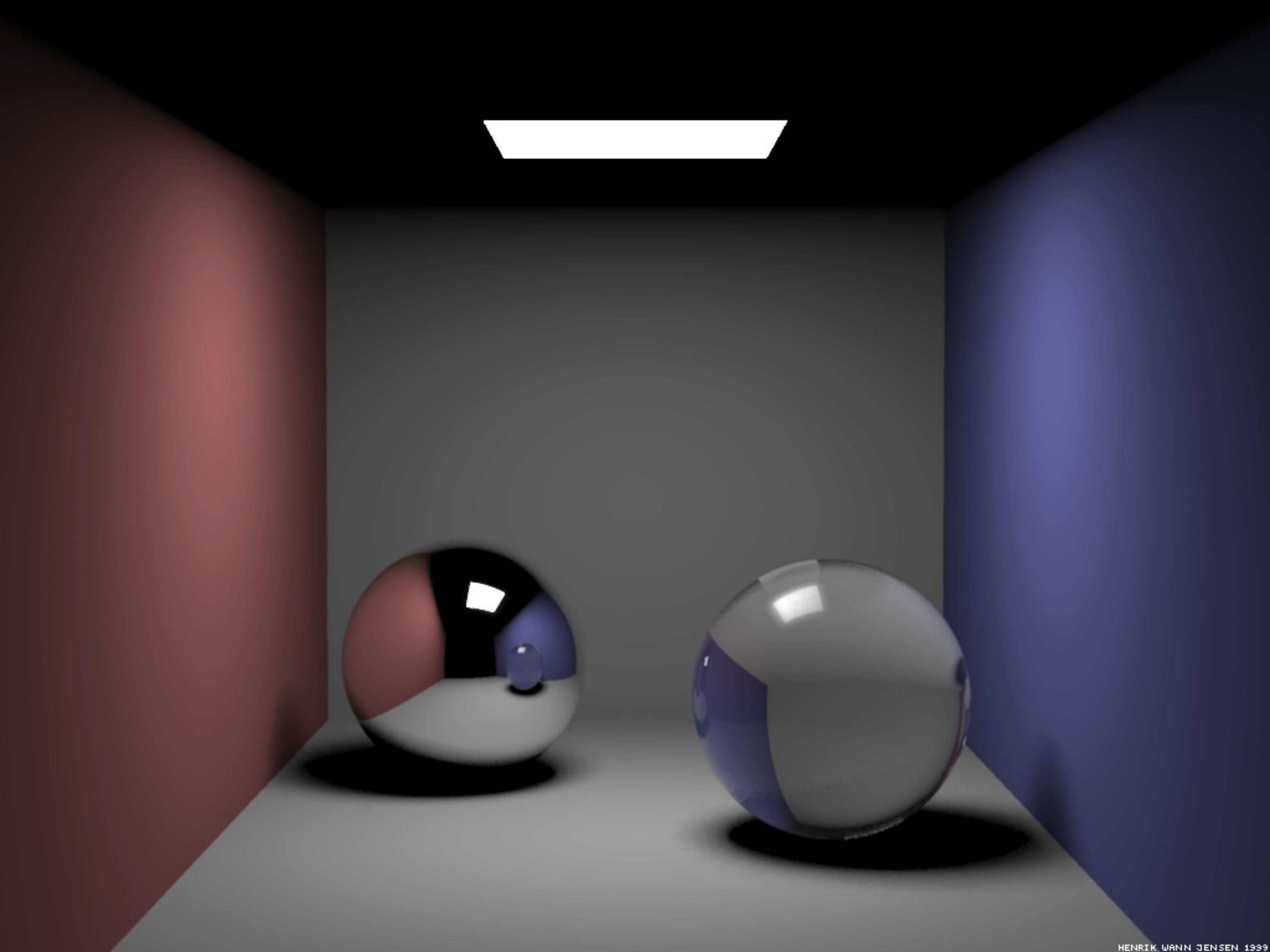
- *Can be done by **photon mapping**.*
- *Shoot light particles (photons) from light sources. They behave like rays.*
- *Store information where they hit surfaces.*
- *Render lighting from the map.*

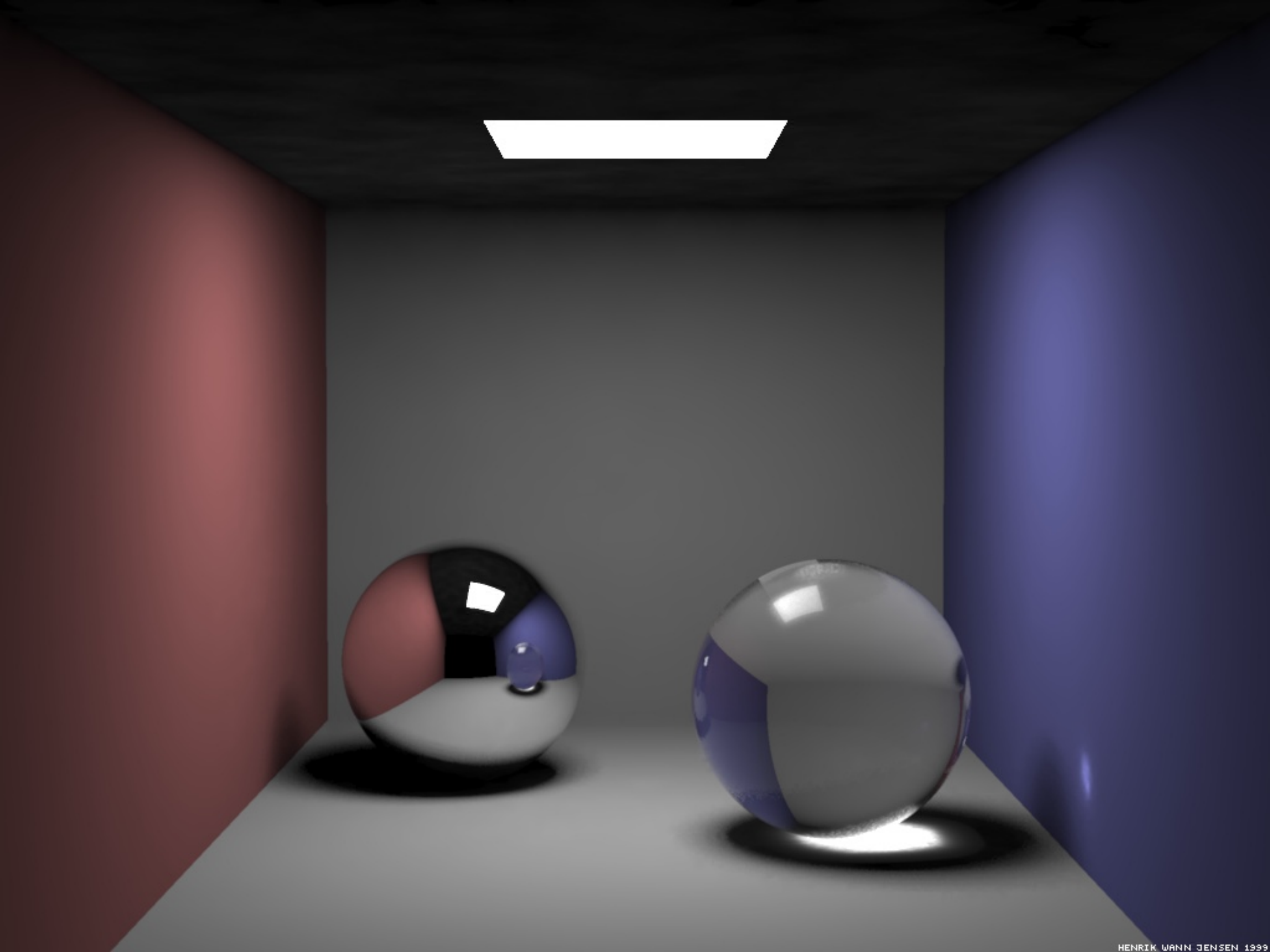


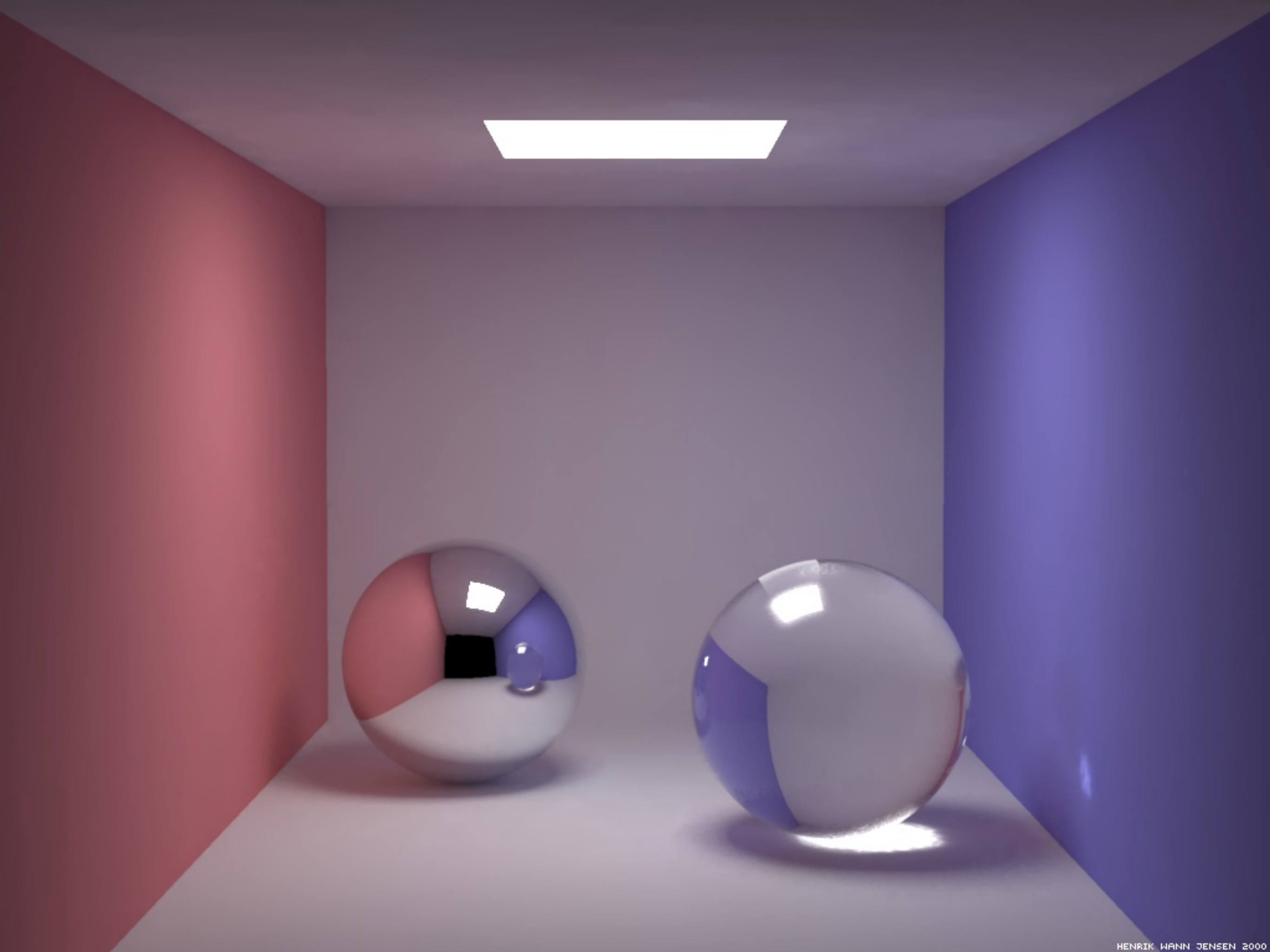
HENRIK WANN JENSEN - 2002











Links

- *More on global illumination*
<http://escience.anu.edu.au/lecture/cg/GlobalIllumination/printNotes.en.html>
- *Our radiosity example used:*
<http://dudka.cz/rrv/gallery?lang=cz>
- *Cornell Box model:*
<http://graphics.ucsd.edu/~henrik/images/cbox.html>